

Vesmírné teleskopy - ten největší a první český

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AOVABVRVONO IN TURNEYOR

A CENTRE * MUNI PELIX

High energy astrophysics group (HEA)



https://hea.physics.muni.cz/



Brief characteristics

Biggest and most sensitive infrared telescope





Infrared radiation

- Longer wavelength than the radiation that we observe
- Shorter wavelength than radio waves



Everything "warm" radiates at infrared wavelengths: atmosphere, Earth, Moon, Sun



James Webb Space Telescope Largest space telescope



Full scale model at NASA Goddard



Infrared radiation

- Longer wavelength than the radiation our eyes "detect"
- Shorter wavelength than radio waves
- astronomer Kris Sellgren (Galactic center observations)



Detection of infrared radiation

- With Earth-bound telescopes we can observe at specific wavebands where the absorption by atmosphere water vapor and other molecules is small
- "Warm" atmosphere is also the source of noise



Infrared space telescope avoids the disturbing effects of the atmosphere

Telescope is also a bit warm, which is still a source of noise → necessary passive and active cooling

Infrared radiation in astronomy

Reasons to construct a big (infrared) space telescope:

- Distant galaxies have their spectrum "redward" shifted the emission peak in the infrared domain
- Infrared radiation penetrates more easily through dense and cool gas and dust clouds (detection of young, forming stars)
- Cooler objects like protoplanetary disks and planets radiate mostly in the infrared domain
- Infrared radiation, when observed from the Earth, is affected by the atmosphere \rightarrow put it into space



• 1946 - Lyman Spitzer created the concept of space telescope (RAND Co.)



Document title: Lyman Spitzer, Jr., "Astronomical Advantages of an Extra-terrestrial Observatory," Project RAND, July 30, 1946.

Source: The RAND Corporation, reprinted with permission.

Priot to World War II, Earth-orbiting telescopes only existed in science fiction stories. The advent of guided missiles by G emany during the war, however, made a few astronomers optimistic that this new rocket technology would soon be able to loft telescopes and other astronomical instruments into space. A mong the balievers, Princeton University's Lyman Spitzer authored a paper for the Douglas Aircraft Company S Project RAND. (the think tank established by the Army Air Corps after World War II) on the scientific benefits of a spacebased telescope The paper became part of a larger 1946 RAND report on the feesbility of developing and launching a scientific spaceraft. Originally dassified, the Spitzer study was unknown to other astronomers for several years. When this ideas became known, many astronomers remained skeptical of the worth of spacebased instruments. Over time, however, astronomers begin to embrace the astronomical sudies Spitzer described in his paper and eventually attributed the H ubble Space T descope's development to Spitzer's efforts.

[no page number] YALE UNIVERSITY OBSERVATORY PROSPECT AND CANNER STREETS NEW HAVEN 11, CONNECTICUT ASTRONOMICAL ADVANTAGES OF AN EXTRA-TERRESTRIAL OBSERVATORY

It has been proposed that rockets be used to accelerate a small mass, containing scientific equipment, up to a speed of 5 miles a second, at which speed the mass could revolve around the earth indefinitely, forming a small satellite. Such a development is certainly not out of the question within the next few decades, in view of the rapid strides already made in rocket research, and the emphasis now being placed on research in this field. The present memorandum points out, in a very preliminary way, the results that might be expected from astronomical measurements made with such a stellite. The discussion is divided into three parts, corresponding to three different assumptions concerning the amount of instrumentation provided. In the first section it is assumed that no telescope is provided; in the second a 10-inch reflector is assumed; in the third section some of the results obtainable with a large reflecting telescope, many feet in diameter, and revolving about the earth solve the terrestrial atmosphere, are briefly sketched. It should be emphasized that this is only a preliminary survey of the scientific advantages that astronomy might gain from such a development. The many practical problems



• Launching of 4 big space telescopes "Great Observatories" between 1990 and 2003

July 30, 1946.

 Compton GRO (1991-2000), Chandra X-ray Observatory (1999-), Hubble Space Telescope (1990-), Spitzer Infrared Telescope (2003-2020)



- 1996 the concept of Next Generation Space Telescope
- 2002 naming after James E. Webb (1906-1992), second administrator of NASA (1961-1968) during the era of the first human space missions Mercury and Gemini
- July 2011 project nearly cancelled
- November 2011 project saved
- 2013-2014 integration of 4 detectors into the

Integrated Science Instrument Module (ISIM)





• **2013-2014** – integration of 4 detectors into

Integrated Science Instrument Module (ISIM)

- <u>Observation modes</u>: imaging, spectroscopy (decomposition into "rainbow"), coronography (direct exoplanet imaging by blocking the stellar light), combination of imaging and spectroscopy (IFU integral field unit), aperture interferometry
- NIRcam, NIRspec, FGS/NIRISS, MIRI





- June 2014 cryogenic test with all the four detectors at the Goddard space center
- 2015 hexagonal mirror segments completed
- **2016** cryogenic test of mirrors and other instruments
- November, 2016 basic telescope construction completed
- December, 2016 anomaly during vibration tests
- 2018 launch postponed until 2020, problems during the propulsion tests and fracture of the antisolar shield
- 2020-2021 delays due to COVID19, rocket Ariane 5







Launch and journey to the point L2

- December 25, 2021 12:20 UTC: rocket Ariane 5 with JWST launched from Kourou, French Guiana
- January 24, 2022 JWST is inserted into the orbit around the L2 point of the Sun-Earth system, 1.5 million kilometers from the Earth (3-4 times more distant than the Moon from the Earth)



Lagrange equilibrium points

 5 points: L1, L2, L3 along the Sun-Earth line (unstable); L4 and L5 at the peaks of the equilateral triangle (stable)



Joseph-Louis Lagrange



(1736-1813)

3-body problem → Langrange points

JWST orbits around the L2 point

- JWST orbits around the L2 point once in every ~6 months
- Orbital corrections every 3 months
- Stable temperature conditions, detectors are passively and actively cooled



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JWST orbits around the L2 point

- JWST orbits around the L2 point once in every ~6 months
- Orbital corrections every 3 months
- <u>Halo orbit -</u> periodic 3D orbit around the Lagrangian points. Top view for SOHO:

1995-12-02	SOHO	
		••
0.000km/s	119,133km	



Robert Farquhar (1932-2015) "Father of halo orbits"

Telescope construction



Primárne zrkadlo

- Overall diameter 6.5 meters, composed from 18 hexagonal segments made of beryllium with the gold coat
- Golden layer is only 100 nm thick altogether 48 grams (golf ball has 46 grams)
- Beryllium is a rare metal, it is light as well as flexible
- Each segment has only 20 kg and it is about 1 mm thick



Primary (segmented) mirror

 Development Guido Horn D'Arturo (Bologna, 1952) and Jerry Nelson (Lawrence Berkeley National Laboratory)







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Antisolar shield

• Big as a tennis court: 14 x 21 meters

Cross-Section of Webb's Five-Layer Sunshield

- Five thin layers with vacuum in between good insulation properties
- Material: kapton E covered with the layer of aluminum and silicon





Optical system of JWST



Image of a star provided by the optical system





Scientific focus

 Origin of stars and planets, their evolution in the Universe across the galaxy and time



Scientific focus

- JWST has been studying formation and evolution of galaxies from 180 million years after the Big Bang (hot and dense origin of our Universe)
- First stars formed only 100 million years after the Big Bang
- From ~100 million years (z=30) till ~914 million years (z=6), the neutral hydrogen was reionized
- In our Galaxy, we will study star formation, exoplanets, and the Galactic center



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Webb's First Deep Field SMACS J0723.3-7327

Galaxy cluster - 4.6 billion light years distant, some "red" galaxies ~13.1 Gly



MIRI



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NIRSpec Microshutter Array Spectroscopy

Stephan's quintet

4 colliding galaxies - 290 million light years from the Earth



Carina nebula

8500 light years from the Earth, star-formation region (NGC3324)





MIRI

NIRCam

Southern Ring Nebula (NGC 3132)

Gas and dust expelled by a dying star (2500 light years from the Earth)



NIRCam



WASP-96b

Analysis of the exoplanet atmosphere spectrum, detection of water and clouds (1120 ly)



NIRISS

WASP-96b

Analysis of the exoplanet atmosphere spectrum, detection of water and clouds (1120 ly)



Cartwheel galaxy ~ 500 milion light years



NIRCam/MIRI

NIRCam

MIRI

Jupiter with moons and the ring



NIRCam

NIRCam

NIRCam

M74 - Fanthom galaxy - 32 million light years from the Earth



NGC1365 - Great Barred Spiral Galaxy - 56 million light years



Visible light (DESCam)

Tarantula nebula (30 Doradus) - HII region in Giant Magellanic cloud - 160 000 light years from the Earth



NIRCam

First detection of carbon dioxide in the exoplanet atmosphere WASP-39b: 700 light years distant, temperature 900 degrees Celsius, 4-day orbit around the star



Neptune with rings and moons





NIRCam



Exoplanet imaging by blocking the stellar light



Protostar - forming star in the dark cloud L1527 at 460 light years in the Taurus star-forming region



Potential biosignature?



- Detection of methane, carbon dioxide, and traces of dimetyl sulfide in the atmosphere of the transiting exoplanet K2-18b
- Subneptune, Hycean (hydrogen+ocean) planet with potentially water ocean and hydrogen atmosphere (lack of ammonia)
- 120 light years from the Earth





Potential biosignature?

• Detection of CO2 on Europa's surface



NIRSpec/IFU

Oldest and most distant single star - <u>Earendel (1 billion years after the</u> Big Bang). B-type star, temperature 2x Solar, 10⁶ Solar luminosities



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Oldest and most distant single star - <u>Earendel</u> Extreme magnification by 4000, likely a binary with a cooler companion



Anniversary image

Nearest star-forming region at 390 ly





NIRCam

Herbig-Haro object 46/47 at 1470 ly Jet from a forming binary star colliding with a dark nebula (Bok nebule)





New Technology Telescope (ESO)

Near-infrared image (JWST)

Herbig-Haro object 46/47 at 1470 ly Jet from a forming binary star colliding with a dark nebula (Bok nebule)



Herbig-Haro object 46/47 at 1470 ly Mysterious object in the background



Herbig-Haro object 46/47 at 1470 ly Mysterious object in the background



Herbig-Haro object 46/47 at 1470 ly Mysterious object in the background



Mysterious objects in Orion



M42: 1400 light years distant





JuMBOs - Jupiter-Mass Binary Objects

Planned/ongoing observations

Galactic center - our project of investigating young stars in the vicinity of the supermassive black hole (Sgr A*)



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K/L-band, 2.2/3.6 µm



QUVIK Small space telescope



QUVIK Small space telescope





https://quvik.cz/

UV astronomy

- Two UV bands (near and far UV)
- Main research focus: kilonovae (mergers of 2 neutron stars)
- Other areas
 Stars (supernovae, hot stars)
 Active galactic nuclei
 (tidal disruption events)
- Planned launch: 2028



Ďakujem za pozornosť!!!



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